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# Long-range stiffness of spectral fluctuations in integrable scale-invariant systems 

T H Seligman $\dagger$ and J J M Verbaarschot $\ddagger$<br>$\dagger$ Instituto de Fisica, University of Mexico in Cuernavaca (UNAM), AP 01000 Mexico DF, Mexico<br>$\ddagger$ Department of Physics, Loomis Laboratory, University of Illinois at Urbana-Champaign, 1110 West Green Street, Urbana, 61801 Illinois, USA

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#### Abstract

In this paper we study the semiclassical limit for the $\Delta_{3}$ statistic of integrable systems which have a homogeneous polynomial as potential. These systems possess a scale invariance which provides us with the energy dependence of the statistic using the general expressions given by Berry. We also obtain the functional dependence of the universal part of the $\Delta_{3}$ statistic both for integrable and ergodic systems. We use a stationary phase approximation to evaluate the semiclassical limit of the level density when the dimension of the system is larger than one. The investigation of its validity leads to the introduction of a generalised perimeter term in the average level density. The fluctuating part of the semiclassical level density yields the semiclassical limit of the $\Delta_{3}$ statistic, which is compared numerically to results obtained for actual spectra. We find that the semiclassical approximation is excellent provided the perimeter term is taken into account exactly. We also study the dimensional dependence of $\Delta_{3}$. For energy levels around the $N$ th level above the ground state the position of the kink in the $\Delta_{3}$ statistic is essentially proportional to $N^{(d-1) / d}$.


## 1. Introduction

Recently quantum spectra of classically chaotic systems have received a considerable amount of attention (see Bohigas and Giannoni (1984) for a review). Both analytical and numerical evidence has been found that the spectral fluctuations of completely chaotic systems can be described by the invariant random matrix ensembles. Most recently, Berry (1985) was able to give a semiclassical derivation of the correct logarithmic dependence of the $\Delta_{3}$ statistic. This statistic is a function of $L$, the length of an interval along the spectrum in units of average level spacings. It measures the fluctuations of the number of quantum levels in intervals of length $L$. The transition region between completely chaotic and integrable systems has been subjected to several studies (Yukawa 1985, Ishikawa and Yukawa 1985, Seligman et al 1984, 1985) but has not yet been fully understood. However, there is also evidence that in this case random matrix theory plays an important role (Seligman et al 1985). In this context a problem arose as, close to the integrable limit, the non-universal properties of the level correlation function are very pronounced. In particular we refer to the fact that the $\Delta_{3}(L)$ statistic no longer saturates after a certain value of $L$. This so-called 'kink' was first reported by Zirnbauer and the present authors (Seligman et al 1984). After that it was studied by Casati et al (1985) for the square well. In the same study they showed that even the nearest-neighbour spacing distribution of their Hamiltonian is non-generic (see also Seligman and Verbaarschot 1986a).

Recently, Berry (1985) explained the phenomenon of the 'kink' by using the semiclassical limit of the fluctuating part of the level density. He was able to give a quantitative explanation of the data of Casati et al (1985). However, his theory has not been tested for systems other than the square well. It is the objective of this paper to perform this task. We will concentrate on the class of separable homogeneous potentials. This class has the advantage that all calculations can be performed explicitly. We take this opportunity to make a careful analysis of the relevant stationary phase approximations.

Since actual physical systems usually involve many degrees of freedom we have also been interested in the dimensional dependence of the non-universal properties of the $\Delta_{3}$ statistic. Intuitively it is clear that the position of the 'kink' will move to larger values of $L$ when we increase the dimensionality of the system. We shall also investigate this surmise for the aforementioned systems.

The organisation of this paper is as follows. In $\S 2$ we define the model and use general scaling arguments to derive the energy dependence of the $\Delta_{3}$ statistic. In § 3 we evaluate the semiclassical limit of the fluctuating part of the level density. The stationary phase approximation is discussed and a generalised perimeter correction to the average level density is derived. In § 4 we derive an explicit expression for the $\Delta_{3}$ statistic. Numerical results are presented in § 5 and concluding remarks are made in $\S 6$.

## 2. Scaling properties of the $\Delta_{3}$ statistic

We consider integrable systems with a homogeneous polynomial as potential. In a $d$-dimensional space the Hamiltonian is given by

$$
\begin{equation*}
H=\frac{1}{2} \sum_{i=1}^{d} p_{i}^{2}+V^{(2 m)}\left(x_{1}, \ldots, x_{d}\right) \tag{2.1a}
\end{equation*}
$$

where

$$
\begin{equation*}
V^{(2 m)}\left(x_{1}, \ldots, x_{d}\right)=\sum_{i=1}^{d} c_{i} x_{i}^{2 m} \tag{2.1b}
\end{equation*}
$$

Hamiltonians of the form (2.1) have scaling properties. Solutions of the Hamilton equations of motion at different energies are related by a simple scale transformation of the coordinates and the time (Landau and Lifshitz 1967). For the wкв approximation to the eigenvalues $E$ this implies that

$$
\begin{equation*}
E=\sum_{i=1}^{d} \gamma_{i}\left(n_{i}+\alpha_{i}\right)^{p} \hbar^{p} \quad p=2 m /(m-1) . \tag{2.2}
\end{equation*}
$$

Here, the $n_{i}$ are integers larger than or equal to zero and the $\gamma_{i}$ and $\alpha_{i}$ are constants to be discussed below. Equation (2.2) follows immediately by expressing the Hamiltonian in action-angle variables. The actions along one period of a periodic trajectory are

$$
\begin{equation*}
I_{i}=\frac{1}{2 \pi} \oint p_{i} \mathrm{~d} x_{i}=a_{i} E_{i}^{1 / 2+1 / 2 m} \tag{2.3a}
\end{equation*}
$$

The constants $a_{i}$ are given by

$$
\begin{equation*}
a_{i}=\frac{\sqrt{2}}{\pi m} c_{i}^{-1 / 2 m} B\left(\frac{3}{2}, \frac{1}{2 m}\right) \tag{2.3b}
\end{equation*}
$$

where $B(x, y)$ is the beta function. In terms of the variables $I_{i}$ the Hamiltonian can
be written as

$$
\begin{equation*}
H=\sum_{i=1}^{d}\left(\frac{I_{i}}{a_{i}}\right)^{2 m /(m+1)} \tag{2.4}
\end{equation*}
$$

In (2.2) the values of the $\alpha_{i}$ are given by

$$
\begin{array}{ll}
\alpha_{i}=0.5 & \text { for } 1<p<\infty \\
\alpha_{i}=0 & \text { for } p=\infty \text { (von Neumann boundary conditions) }  \tag{2.5}\\
\alpha_{i}=1.0 & \text { for } p=\infty \text { (Dirichlet boundary conditions). }
\end{array}
$$

From (2.4) one immediately deduces the wкв approximation given in (2.2) (the constants $\gamma_{i}$ are given by $\gamma_{i}=a_{i}^{-p}$ ). At this point we want to remark that for Hamiltonians with a discrete symmetry the value of $\alpha_{i}$ is different from (2.5) when we consider only levels of one symmetry class. For instance, the Hamiltonian (2.1) is invariant under the parity operation. The odd and even parity states will be considered separately. In a one-dimensional system the even parity levels are given by

$$
\begin{equation*}
E_{n}=\gamma\left(2 n+\frac{1}{2}\right)^{p}=\gamma 2^{p}\left(n+\frac{1}{4}\right)^{p} \tag{2.6}
\end{equation*}
$$

and the odd parity levels by

$$
\begin{equation*}
E_{n}=\gamma\left(2 n+1+\frac{1}{2}\right)^{p}=\gamma 2^{p}\left(n+\frac{3}{4}\right)^{p} . \tag{2.7}
\end{equation*}
$$

We have used the fact that in a one-dimensional potential of the form (2.1) the levels are ordered according to the number of nodes in the wavefunction. The value of $\alpha_{i}$ is of importance for the discussion in $\S 3$.

The accuracy of the wкв approximation has been studied numerically by diagonalising Hamiltonians of the form (2.1) in a harmonic oscillator basis. As an illustration we show in the second column of table 1 some eigenvalues corresponding to positive parity states of an $x^{4}$ potential (the index of the eigenvalues is given in the first column). In the third column we give the wКв approximation calculated from $c\left(n+\frac{1}{4}\right)^{4 / 3}(c=$ 1.1720340504 ). The agreement is very good. Analytical estimates have been given by Hioe et al (1978). They found that the deviation from the wкв result is of order $n^{-2}$ in the one-dimensional case. Therefore, to leading order in $n^{-1}$ these deviations do not contribute to the correlations of the eigenvalues in higher-dimensional systems. This will become clear in § 3. In conclusion, we can use (2.2) as a starting point for the study of correlations of eigenvalues of Hamiltonians (2.1).

We will restrict our study to correlations that extend over a large but in the limit $\hbar \rightarrow 0$ finite number of eigenvalues. The appropriate measure for these correlations is

Table 1. Comparison of exact eigenvalues and their wKB approximation for a onedimensional $x^{4}$ potential.

| $n$ | Exact value | WKB value |
| ---: | :---: | :---: |
| 0 | 0.185 | 0.226 |
| 1 | 1.578 | 1.587 |
| 2 | 3.456 | 3.461 |
| 5 | 10.694 | 10.697 |
| 10 | 26.096 | 26.098 |
| 25 | 86.820 | 86.822 |
| 50 | 217.331 | 217.332 |
| 100 | 538.577 | 538.577 |

the $\Delta_{3}$ statistic defined by

$$
\begin{equation*}
\Delta_{3}(L)=\min _{A, B} \frac{\bar{\rho}}{L} \int_{-L / 2 \bar{\rho}}^{L / 2 \bar{\rho}} \mathrm{~d} \varepsilon(N(E+\varepsilon)-A-B \varepsilon)^{2} \tag{2.8}
\end{equation*}
$$

Here, $N(E)$ is the integrated level density (a staircase function) and $\bar{\rho}$ is the average level density. The minimisation is over $A$ and $B$. For the evaluation of $\Delta_{3}$ we will use the semiclassical approximation to the level density with inclusion of the lowest order quantum fluctuations. According to the work of Gutzwiller (1967) $\rho(\varepsilon)$ is given by (see also Berry and Mount 1972)

$$
\begin{equation*}
\rho(\varepsilon)=h^{-d} \int \delta(E-H) \mathrm{d}^{d} x \mathrm{~d}^{d} p+\hbar^{-(d+1) / 2} \sum_{j} A_{j}(E) \exp \left(\mathrm{i} S_{j}(E) / \hbar\right) \tag{2.9a}
\end{equation*}
$$

for integrable systems and

$$
\begin{equation*}
\rho(\varepsilon)=h^{-d} \int \delta(E-H) \mathrm{d}^{d} x \mathrm{~d}^{d} p+\hbar^{-1} \sum_{j} A_{j}(E) \exp \left(\mathrm{i} S_{j}(E) / \hbar\right) \tag{2.9b}
\end{equation*}
$$

for ergodic systems. The sum is over all closed periodic trajectories with energy $E$. $S_{j}(E)$ is the action along one such trajectory and $A_{j}(E)$ are coefficients that can be derived from the properties of the classical trajectories. Note that the physical dimension of the amplitudes $A_{j}(E)$ is different in the integrable and ergodic cases. Since a number theoretic derivation of (2.9a) will be given in $\S 3$, we do not quote the explicit expressions for $S_{j}(E)$ and $A_{j}(E)$. For the interested reader we refer to the work of Gutzwiller (1967, 1970, 1971) and Berry and Tabor (1977). By using the scaling properties of the Hamiltonian (2.1) we will derive a result for the energy dependence of $\Delta_{3}(L)$. For the study of the eigenvalue correlations in an interval around $E$ containing a finite number of eigenvalues we can make the following linear approximation to the action:

$$
\begin{equation*}
S_{j}(E+\varepsilon)=S_{j}(E)+\varepsilon \mathrm{d} S_{j} / \mathrm{d} E . \tag{2.10}
\end{equation*}
$$

In the limit $h \rightarrow 0$ this approximation is exact. In this limit the energy dependence of the smooth functions $A_{j}(E)$ is irrelevant (see Berry 1985). By changing to dimensionless integration variables we can write $\Delta_{3}(L)$ as

$$
\begin{equation*}
\Delta_{3}(L)=\min _{A, B} \int_{-1}^{1} \mathrm{~d} \varepsilon\left[N\left(E+\frac{\varepsilon L}{2 \bar{\rho}}\right)-A-B \varepsilon\right]^{2} . \tag{2.11}
\end{equation*}
$$

Due to the minimisation the constant multiplying $B$ can be omitted. Only the oscillating part of $N(E)$ contributes to $\Delta_{3}$. Since $\varepsilon$ is dimensionless $\Delta_{3}(L)$ is a function of $(L / 2 \bar{\rho} \hbar) \mathrm{d} S_{j} / \mathrm{d} E$. For scaling systems the energy dependence can be derived from dimensional arguments only. For the relevant quantities we find the following dependence on $\hbar$ :

$$
\begin{array}{cc}
\frac{L}{2 \bar{\rho} \hbar} \frac{\mathrm{~d} S_{j}}{\mathrm{~d} E} \sim[S]^{1-d} \hbar^{d-1} & \\
\hbar^{1-d}\left(\frac{\mathrm{~d} S_{j}}{\mathrm{~d} E}\right)^{-2} A_{j}^{2} \sim[S]^{d-1} \hbar^{1-d} & \text { for integrable systems } \\
\hbar^{0}\left(\frac{\mathrm{~d} S_{j}}{\mathrm{~d} E}\right)^{-2} A_{j}^{2} \sim[S]^{0} \hbar^{0} & \text { for ergodic systems } \\
N(E) \sim[S]^{d} \hbar^{-d} . & \tag{2.12d}
\end{array}
$$

[ $S$ ] denotes the dimension of an action. Putting everything together results in

$$
\begin{equation*}
\Delta_{3}(L)=N(E)^{(d-1) / d} f\left(L(N(E))^{(1-d) / d}\right) \tag{2.13a}
\end{equation*}
$$

for integrable systems and

$$
\begin{equation*}
\Delta_{3}(L)=g\left(L(N(E))^{(1-d) / d}\right) \tag{2.13b}
\end{equation*}
$$

for ergodic systems, where $f$ and $g$ are functions that do not depend on $\hbar, L$ and $N(E)$. This energy dependence in (2.13b) is consistent with the saturation value of the $\Delta_{3}$ statistic derived by Casati et al (1985) for the square well. The generalisation of their argument to an arbitrary scaling potential yields the same asymptotic dependence. From ( $2.13 a$ ) we also conclude that the universal part of the $\Delta_{3}$ statistic, i.e. the part that does not depend on $N(E)$, has to be linear in the integrable case. In the ergodic case we consider the difference $\Delta_{3}(L)-\Delta_{3}(1)$. It is clear that the universal part of this statistic can only have a logarithmic dependence on $L$. Berry (1985) derived the asymptotic limit of $\Delta_{3}$ by using a sum rule for the amplitudes $A_{j}$ derived by Hannay and Ozorio de Almeida (1984). Up to constants this sum rule can be derived by scaling arguments for systems with homogeneous polynomials as a potential (Seligman and Verbaarschot 1986b).

From (2.12) it is clear that an expansion in powers of $\hbar$ is also an expansion in inverse powers of $N(E)$.

## 3. Semiclassical limit of the level density

In § 2 we have explained that the eigenvalues of Hamiltonians of the form (2.1) are given by the wкв approximation (2.2). Berry and Tabor $(1976,1977)$ have derived general formulae for the semiclassical level density of integrable systems. Since for the eigenvalues (2.2) all calculations can be carried out explicitly without loss of the essential features of the general case we present a derivation of the semiclassical level density. For simplicity we consider the two-dimensional case. The generalisation to arbitrary dimensions is immediate.

The level density following from (2.2) is given by (from now on we put $\hbar=1$ )

$$
\begin{equation*}
\rho(\varepsilon)=\sum_{n_{i}=0}^{\infty} \delta\left[E-\gamma_{1}\left(n_{1}+\alpha_{1}\right)^{p}-\gamma_{2}\left(n_{2}+\alpha_{2}\right)^{p}\right] . \tag{3.1}
\end{equation*}
$$

We use the Poisson summation formula to write the level density as (see Berry and Tabor 1976)
$\rho(E)=\int_{0}^{\infty} \int_{0}^{\infty} \mathrm{d} n_{1} \mathrm{~d} n_{2} \sum_{M_{1} M_{2}} \exp \left\{2 \pi \mathrm{i}\left[M_{1}\left(n_{1}-\alpha_{1}\right)+M_{2}\left(n_{2}-\alpha_{2}\right)\right]\right\} \delta\left(E-\gamma_{1} n_{1}^{p}-\gamma_{2} n_{2}^{p}\right)$.

The $\delta$ function can be evaluated by transforming the $n_{i}$ variables to 'polar' coordinates

$$
\begin{equation*}
n_{1}^{p} \gamma_{1}=r \cos ^{2} \varphi \quad n_{2}^{p} \gamma_{2}=r \sin ^{2} \varphi \tag{3.3}
\end{equation*}
$$

The Jacobian $J$ of this transformation is given by

$$
\begin{equation*}
J=2\left(\gamma_{1} \gamma_{2}\right)^{-1 / p} p^{-2}(r \cos \varphi \sin \varphi)^{2 / p-1} \tag{3.4}
\end{equation*}
$$

We note that for $0<p<2$ this Jacobian is zero for $\varphi$ equal to zero and $\pi / 2$. The $r$ integration can be carried out. This yields for the level density

$$
\begin{align*}
& \rho(E)=\frac{2}{p^{2}} E^{2 / p-1}\left(\gamma_{1} \gamma_{2}\right)^{-1 / p} \int_{0}^{\pi / 2} \mathrm{~d} \varphi \sum_{M_{1}}\left(\frac{\sin 2 \varphi}{2}\right)^{2 / p-1} \\
& \quad \times \exp \left\{2 \pi \mathrm{i}\left[M_{1}\left(\frac{E}{\gamma_{1}}\right)^{1 / p} \cos ^{2 / p} \varphi+M_{2}\left(\frac{E}{\gamma_{2}}\right)^{1 / p} \sin ^{2 / p} \varphi-\alpha_{1} M_{1}-\alpha_{2} M_{2}\right]\right\} . \tag{3.5}
\end{align*}
$$

When both $M_{1}$ and $M_{2}$ are different from zero the $\varphi$ integration can be performed by the method of stationary phases. In the limit $E \rightarrow \infty$ the results thus obtained become exact. When either $M_{1}$ or $M_{2}$ equals zero the stationary point is at the boundary of the integration manifold. Just at these points the Jacobian $J$ in (3.4) is equal to zero for $0<p<2$. Moreover, for $0<p<2$ there are two stationary points, one at zero and the other at $\pi / 2$. The second derivative with respect to $\varphi$ at one of these points is infinite for $1<p<2$. (For billiard problems ( $p=2$ ), where most of these problems do not occur, the stationary points at the boundaries have been discussed by Richens and Berry (1981).) This suggests that to make a careful analysis of this case we should not transform the $n_{i}$ variables to 'polar' coordinates but instead use (3.2) as a starting point. We will add the terms with either $M_{1}$ or $M_{2}$ equal to zero to the term with both $M_{1}$ and $M_{2}$ equal to zero. Together they define the average level density $\rho_{\mathrm{SM}}(E)$. We will evaluate these terms exactly. Thereby we circumvent the problems involving the stationary phase approximation. For $\rho_{\mathrm{SM}}$ we obtain

$$
\begin{align*}
\rho_{\mathrm{SM}}(E)=\int_{0}^{\infty} & \int_{0}^{\infty} \mathrm{d} n_{1} \mathrm{~d} n_{2}\left(1+\sum_{M_{1} \neq 0} \exp \left[2 \pi \mathrm{i} M_{1}\left(n_{1}-\alpha_{1}\right)\right]\right. \\
& \left.+\sum_{M_{2} \neq 0} \exp \left[2 \pi \mathrm{i} M_{2}\left(n_{2}-\alpha_{2}\right)\right]\right) \delta\left(E-\gamma_{1} n_{1}^{p}-\gamma_{2} n_{2}^{p}\right) \\
= & \int_{0}^{\infty} \int_{0}^{\infty} \mathrm{d} n_{1} \mathrm{~d} n_{2}\left(\sum_{k_{1}=0}^{\infty} \delta\left(n_{1}-\alpha_{1}-k_{1}\right)\right. \\
& \left.+\sum_{k_{2}=0}^{\infty} \delta\left(n_{2}-\alpha_{2}-k_{2}\right)-1\right) \delta\left(E-\gamma_{1} n_{1}^{p}-\gamma_{2} n_{2}^{p}\right) \tag{3.6}
\end{align*}
$$

The integrations can be carried out easily. This yields

$$
\begin{align*}
& \rho_{\mathrm{SM}}(E)=\frac{1}{p} \gamma_{2}^{-1 / p} \sum_{\substack{k_{1}=0 \\
\gamma_{1}\left(k_{1}+\alpha_{1}\right)^{p}<E}}\left[E-\gamma_{1}\left(k_{1}+\alpha_{1}\right)^{p}\right]^{1 / p-1}+\frac{1}{p} \gamma_{1}^{-1 / p} \\
& \times \sum_{\substack{k_{2}=0 \\
\gamma_{2}\left(k_{2}+\alpha_{2}\right)^{p}<E}}\left[E-\gamma_{2}\left(k_{2}+\alpha_{2}\right)^{p}\right]^{1 / p-1}-\frac{1}{p^{2}}\left(\gamma_{1} \gamma_{2}\right)^{-1 / p} E^{2 / p-1} B\left(\frac{1}{p}, \frac{1}{p}\right) \tag{3.7}
\end{align*}
$$

where $B(x, y)$ is the beta function. When $E \rightarrow \infty \rho_{\mathrm{SM}}(E)$ tends to the Thomas-Fermi approximation for the average level density. However, for the calculation of the correlation function we have to keep all terms in (3.7). We evaluate the corrections in leading order in $E^{-1 / p}$. Since the first two terms in (3.7) are both Riemann sums
that approximate the integral given by the last term, in lowest order in $E^{-1 / p}$ the corrections to the leading order term $p^{-2}\left(\gamma_{1} \gamma_{2}\right)^{-1 / p} E^{2 / p-1} B(1 / p, 1 / p)$ are given by

$$
\begin{equation*}
\frac{1}{2 p} \gamma_{2}^{-1 / p}\left(\frac{1}{2}-\alpha_{1}\right) E^{1 / p-1}+\frac{1}{2 p} \gamma_{1}^{-1 / p}\left(\frac{1}{2}-\alpha_{2}\right) E^{1 / p-1} . \tag{3.8}
\end{equation*}
$$

By inspection of the potential (2.1) one immediately concludes that for $p=2$, (3.8) yields the perimeter term of the semiclassical level density of a square box (see Richens and Berry 1981, Bohigas and Giannoni 1984). One also observes that this term changes sign going from von Neumann boundary conditions ( $\alpha_{i}=0$ ) to Dirichlet boundary conditions ( $\alpha_{i}=1$ ). When the levels are given by the wкв approximation ( $\alpha_{i}=0.5$ ) there is no perimeter term. However, when one considers levels of a definite symmetry class the value of $\alpha_{i}$ is no longer 0.5 (see (2.5)) and one has to take into account the term given in (3.8). Actually, the higher-order terms in (3.7) belong to the fluctuating part of the spectrum. It is because of the problems in calculating these terms by a stationary phase approximation that we interpret $\rho_{\mathrm{SM}}(E)$ as the average level density.

The fluctuating part of the level density is given by the terms in (3.5) for which neither $M_{1}$ nor $M_{2}$ is equal to zero. The $\varphi$ integration will be performed by the method of stationary phases. In this case there are no additional problems. The stationary point is given by

$$
\begin{equation*}
M_{1} \gamma_{1}^{-1 / p} \cos ^{2 / p-2} \varphi=M_{2} \gamma_{2}^{-\boldsymbol{\beta} / p} \sin ^{2 / p-2} \varphi \tag{3.9}
\end{equation*}
$$

and the second derivative of the exponent at this point is equal to

$$
\begin{equation*}
2 \pi \mathrm{i} E^{1 / p} \frac{4(1-p)}{p^{2}}\left[\left(M_{1} \gamma_{1}^{-1 / p}\right)^{p /(p-1)}+\left(M_{2} \gamma_{2}^{-1 / p}\right)^{p /(p-1)}\right]^{(p-1) / p} . \tag{3.10}
\end{equation*}
$$

Using these results we obtain the following stationary phase approximation for the fluctuating part of the level density:

$$
\begin{align*}
\rho_{\mathrm{osc}}(E)= & \frac{1}{p(p-1)^{1 / 2}} E^{3 / 2 p-1} \sum_{\substack{M_{1}, M_{2} \neq 0 \\
\operatorname{sgn}\left(M_{1}\right)=\operatorname{sgn}\left(M_{2}\right)}} \frac{\left(\gamma_{1} \gamma_{2}\right)^{1 / 2(1-p)}\left(M_{1} M_{2}\right)^{(2-p) / 2(p-1)}}{\left[\left(\gamma_{1}^{1 / p}\right)^{p /(p-1)}+\left(M_{2} / \gamma_{2}^{1 / p}\right)^{p^{/(p-1)}}\right]^{(3-p) / 2 p}} \\
& \times \exp \left\{2 \pi \mathrm{i} E^{1 / p}\left[\left(M_{1} / \gamma_{1}^{1 / p}\right)^{p /(p-1)}+\left(M_{2} / \gamma_{2}^{1 / p}\right)^{p /(p-1)}\right]^{(p-1) / p}-\frac{1}{4} \pi \mathrm{i} \operatorname{sgn} M_{1}\right\} . \tag{3.11}
\end{align*}
$$

This result will be the starting point of the calculation of the $\Delta_{3}$ statistic.
We do not give the derivation for the $d$-dimensional case here but just quote the final result:

$$
\begin{align*}
& \rho_{\mathrm{osc}}(E)=\sum_{M_{i} \neq 0} A_{M} \exp \left(\mathrm{i} S_{M}\right) \\
& A_{M}=\left(\prod_{i} \gamma_{i}\right)^{1 / 2(1-p)}\left(\prod_{i} M_{i}\right)^{(2-p) / 2(p-1)} E^{(d+1) / 2 p-1} \frac{(p-1)^{1 / 2}}{p}(p-1)^{-d / 2} \\
& \times\left[\sum_{i=1}^{d}\left(\frac{M_{i}}{\gamma_{i}^{1 / p}}\right)^{p /(p-1)}\right]^{[-(p-1) / 2 p][d /(p-1)-1]} \exp \left(\frac{1}{4} \pi i \beta\right) \\
& S_{M}=2 \pi E^{1 / p}\left[\sum _ { i = 1 } ^ { d } \left(\frac{M_{i}}{\left.\left.\gamma_{i}^{1 / p}\right)^{p /(p-1)}\right]^{(p-1) / p}} .\right.\right. \tag{3.12}
\end{align*}
$$

(For the definition of $\beta$ see Berry and Tabor (1976).) For the interpretation of this equation in terms of periodic trajectories we refer to Berry (1985).

The average number of levels $N(E)$ below energy $E$ is obtained by integrating $\rho_{\mathrm{SM}}$ up to $E$. We will express the energy $E$ in (3.11) and (3.12) in terms of $N(E)$. In doing this the additional terms given by (3.7) give rise to contributions that vanish in leading order in $E^{-1 / p}$. Therefore we only use the Thomas-Fermi approximation $\bar{\rho}$ to the level density to express $E$ in $N(E)$. In $d$ dimensions $\bar{\rho}$ is given by

$$
\begin{align*}
& \bar{\rho}_{d}(E)=E^{d / p-1}\left(\prod_{i} \gamma_{i}\right)^{-1 / p} p^{-d} I_{d, 1 / p-1,1 / p-1}  \tag{3.13}\\
& I_{d, r, s}=\int \mathrm{d} E_{1} \ldots \mathrm{~d} E_{d} \delta\left(E_{1}+\ldots+E_{d}-1\right)\left(\prod_{i=2}^{d} E_{i}\right)^{r} E_{1}^{s}
\end{align*}
$$

The integral $I_{d, r, s}$ can be solved recursively:

$$
\begin{equation*}
I_{d, r, s}=B(r+1, s+1) I_{d-1, r+s+1, s+1} \tag{3.14}
\end{equation*}
$$

## 4. Evaluation of the $\boldsymbol{\Delta}_{\mathbf{3}}$ statistic

In this section we evaluate the semiclassical limit of the $\Delta_{3}$ statistic for the class of potentials defined in (2.2). Our starting point is the expression for $\Delta_{3}$ as given by Berry (1985)

$$
\begin{equation*}
\Delta_{3}(L)=2 \sum_{\boldsymbol{M}}\left(\frac{A_{M}}{T_{M}}\right)^{2} G\left(\frac{L T_{M}}{2 \bar{\rho}}\right) \tag{4.1}
\end{equation*}
$$

The coefficients $A_{M}$ are given in (3.12) and $T_{M}$ is defined by

$$
\begin{equation*}
T_{M}=\mathrm{d} S_{M} / \mathrm{d} E \tag{4.2}
\end{equation*}
$$

The function $G$ is specific for the correlations described by $\Delta_{3}$ and is given by

$$
\begin{equation*}
G(y)=1-\frac{\sin ^{2} y}{y^{2}}-\frac{3(y \cos y-\sin y)^{2}}{y^{4}} \tag{4.3}
\end{equation*}
$$

After insertion of the formulae for $A_{M}$ and $T_{M}$ in (4.1) and expressing the energy $E$ in $N(E)$ we obtain for $\Delta_{3}(L)$

$$
\begin{align*}
& \Delta_{3}(L)=\frac{2^{d-2}}{\pi^{2}}\left(\frac{p}{2(p-1)}\right)^{d-1}\left(\frac{d}{p I_{d, 1 / p-1,1 / p-1}}\right)^{(d-1) / d}\left(\prod_{i} \gamma_{i}\right)^{-1 /(p-1)+(d-1) / p d}(N(E))^{(d-1) / d} \\
& \times \sum_{M_{i}>0}\left(\prod_{i} M_{i}\right)^{(2-p) /(p-1)}\left[\sum_{i=1}^{d}\left(\frac{M_{i}}{\gamma_{i}^{1 / p}}\right)^{p /(p-1)}\right]^{[(p-1) / p][d /(1-p)-1]} \\
& \times G\left(L \pi(N(E))^{(1-d) / d}\left(\prod_{i} \gamma_{i}\right)^{1 / p d}\right. \\
&\left.\times\left(\frac{p}{d}\right)^{(d-1) / d} I_{d, 1 / p-1,1 / p-1}^{-1 / d}\left[\sum_{i=1}^{d}\left(\frac{M_{i}}{\gamma_{i}^{1 / p}}\right)^{p /(p-1)}\right]^{(p-1) / p}\right) \tag{4.4}
\end{align*}
$$

For $d=1$ there are no fluctuating contributions to the level density and the $\Delta_{3}$ statistic is equal to its minimum value $\left(\frac{1}{12}\right)$. For $d=2$ and $p=2$ (4.4) reduces to the expression for the square well given by Berry (1985). However, there is one difference. We have included the terms with either $M_{1}$ or $M_{2}$ equal to zero in the average level density. Up to the order in $E^{-1 / p}$ we are considering, the result of Berry (1985) differs because
we have treated the perimeter term exactly. Berry evaluated the fluctuating part of the perimeter term via a stationary phase approximation. The stationary points lie on the boundary of the integration manifold. This yields an additional factor of 0.5 for $A_{M}$ and a factor 0.25 for $\boldsymbol{A}_{\boldsymbol{M}}^{2}$.

For large values of $y$ the function $G(y)=1$. Therefore, the value of $\Delta_{3}(L)$ saturates when $L \rightarrow \infty$. From (4.4) it follows that the energy dependence of the saturation value is given by

$$
\begin{equation*}
\Delta_{3}(\infty) \sim(N(E))^{d /(d-1)} \tag{4.5}
\end{equation*}
$$

This result was already derived in § 2 by using scaling arguments. For the square well it has been given by Casati et al (1985).

## 5. Numerical results

In this section we discuss the numerical results. We calculate the $\Delta_{3}$ statistic for a sequence of eigenvalues generated by the WKB approximation given in (2.2) and compare it to the analytical result given in (4.4). In the case of the wкb eigenvalues we unfold the spectrum by means of the semiclassical level density (3.7) (i.e. the spectrum is transformed to a spectrum with average level spacing equal to one but with the same level fluctuations). We study three different potentials of the form (2.1), the homogeneous $x^{4}$ polynomial, the homogeneous $x^{6}$ polynomial and the square well. We consider different energies and different values of the parameters $\gamma_{i}$. In the case of the $x^{4}$ and $x^{6}$ potentials the value of $\alpha_{i}$ is taken to be equal to 0.25 (even parity states). For the square well we take $\alpha_{i}=1.0$.

In figure 1 we present results for the 500 th eigenvalue up to the 1500 th eigenvalue. The dots represent the values of the $\Delta_{3}$ statistic for the wкв eigenvalues and the full curve corresponds to the analytical expression (4.4) calculated for the 1000th eigenvalue. Going from the top to the bottom the ratio $\gamma_{1} / \gamma_{2}$ is equal to $\frac{1}{2}(\sqrt{5}+1), \pi, \Gamma\left(\frac{1}{4}\right)$. The error in the $\Delta_{3}$ statistic is calculated from its variation over the spectrum. For small values of $L$ the error is much smaller than for large values of $L$. The reason is twofold. In the first place the error is reduced by the spectral averaging. In the second place, for large values of $L$ the error has an important systematic component due to the fact that the saturation value of $\Delta_{3}$ changes by a factor $\sqrt{3}$ going from the 500 th to the 1000 th eigenvalue. Since the saturation value of $\Delta_{3}$ is approximately symmetric around the centre of the part of the spectrum under consideration the actual error is less than the error due to the non-stationarity. The large deviation in figure $1(a)\left(x^{4}\right)$ is a consquence only of the large deviations for the 1000th eigenvalue up to the 1500 th eigenvalue. For the 500 th eigenvalue up to the 1000 th eigenvalue and for the 1500 th eigenvalue up the 2000th eigenvalue we do not find any deviations.

In figure 2 we present results for the 4500 th eigenvalue up to the 5500 th eigenvalue. The potentials (also the values of $\gamma_{i}$ and $\alpha_{i}$ ) are the same as in figure 1. Now the saturation value of $\Delta_{3}$ only changes by a factor $(11 / 9)^{1 / 2}$. Therefore our errors are mainly of statistical origin. Indeed, we find that the fluctuations of the dots around the full curve are of the order of the error bars. Moreover the errors are slightly smaller that in the $E=1000$ case (note that the scale on the $Y$ axis differs by a factor 2 ). All curves in figures 1 and 2 look very similar. However, when we take the ratio of $\gamma_{1}$ and $\gamma_{2}$ very different from 1 we find different shapes for the $\Delta_{3}$ statistic. Since this degenerate case is of little physical interest we will not discuss it any further.


Figure 1. Comparison of the $\Delta_{3}$ statistic calculated from the WKB eigenvalues (dots with error bars) and the analytical expression given in (4.4) (full curve) for the 500th eigenvalue up to the 1500 th eigenvalue. The form of the potential is denoted in the figure ( sw refers to square well). The ratio of $\gamma_{1} / \gamma_{2}$ in $(a),(b),(c)$ is given by $\frac{1}{2}(\sqrt{5}+1), \pi, \Gamma\left(\frac{1}{4}\right)$, respectively.

In figure 3 we show the effect of not taking into account the fluctuating part of the perimeter term in the average level density used for unfolding the spectrum. The parameters of the potential are equal to those of figure $1(b)$. In figure $3(a)$ the values of $\alpha_{1}$ and $\alpha_{2}$ are equal to 0.25 and in figure $3(b) \alpha_{1}$ is equal to 0.25 and $\alpha_{2}$ is equal to 0.75 . There are much larger deviations in the asymptotic region than in the corresponding figure in figure 1.

The dimensional dependence of $\Delta_{3}$ is shown in figure 4. We consider the case of a homogeneous $x^{4}$ potential and calculate the expression given in (4.4) for $d=2, d=3$ and $d=4$. The constants $\gamma_{i}$ are taken equal to 1.0 . The figure clearly shows that the value of $L$ for which the $\Delta_{3}$ statistic saturates moves up with dimension. The dimensional dependence is consistent with the scaling relation (2.13).

## 6. Conclusions

In this paper we have studied the $\Delta_{3}$ statistic for systems with a potential given by a separable homogeneous polynomial. We have applied the analysis of Berry (1985) to this class of potentials and found that the $\Delta_{3}$ statistic is reproduced by the analytical


Figure 2. Comparison of the $\Delta_{3}$ statistic calculated from the wKb eigenvalues and the analytical expression given in (4.4) for the 4500 th eigenvalue up to the 5500 th eigenvalue. The parameters of the potential are the same as in the corresponding part of figure 1. For further explanation see figure 1 .
results. The dimensional and energy dependence of the $\Delta_{3}$ statistic could be derived by invoking only the scaling properties of the potential. The results thus obtained were confirmed by the detailed analysis. In order to be able to unfold the spectra given by the wKB approximation (which was found to be excellent) accurately, we had to take into account a generalised perimeter term. This term appeared to have important contributions when the value of the zero-point fluctuations was different from 0.5. After taking into account this term the values of the $\Delta_{3}$ statistic for the wKB eigenvalues could be reproduced satisfactorily. The dimensional dependence of the 'kink' in the $\Delta_{3}$ statistic behaved according to a simple scaling relation. In higher dimensions the 'kink' moves to larger values of $L$. The functional dependence of the universal part of $\Delta_{3}(L)$ could be explained by scaling arguments both for the integrable and the ergodic cases.

A number of practical consequences ensue from these results. First, the 'kink' is irrelevant in nuclear statistics due to the large dimensionality of the system. Second, the 'kink' constitutes a deviation from Poisson statistics ubiquitous for integrable systems, though its location shifts along the spectrum and also depends on the Hamiltonian. Third, the exact availability of the long-range behaviour of the $\Delta_{3}$ statistic for some systems may allow us to extract information from short low-lying spectra.


Figure 3. The $\Delta_{3}$ statistic of the 500 th eigenvalue up the 1500 th WKB eigenvalue obtained without taking into account the perimeter term. The ratio of $\gamma_{1} / \gamma_{2}$ is equal to $\pi$. In ( $a$ ) and (b) the values of the $\alpha_{i}$ are equal to $(0.25,0.25)$ and $(0.25,0.75)$, respectively. For further explanation see figure 1 .


Figure 4. The dimensional dependence of the $\Delta_{3}$ statistic. The full curves show the analytical expression (4.4) calulated for $d=2,3$ and 4 (going from the bottom to the top). The results have been computed for the 100th eigenvalue and the value of $\gamma_{1}$ is equal to 1.0 .

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